

PAPER • OPEN ACCESS

Improvements in the simulation code of the SOX experiment

To cite this article: A. Caminata *et al* 2017 *J. Phys.: Conf. Ser.* **888** 012145

View the [article online](#) for updates and enhancements.

Related content

- [Air Pollution Simulation based on different seasons](#)
Muhaimin
- [Implementation of Hydrodynamic Simulation Code in Shock Experiment Design for Alkali Metals](#)
A L Coleman, R Briggs, M G Gorman et al.
- [Effects of uncertainties in simulations of extragalactic UHECR propagation using CRPropa and SimProp](#)
R. Alves Batista, D. Boncioli, A. di Matteo et al.

Improvements in the simulation code of the SOX experiment

A. Caminata¹, M. Agostini², K. Altenmüller³, S. Appel³,
 V. Atroshchenko⁴, G. Bellini⁵, J. Benziger⁶, D. Bick⁷, G. Bonfini⁸,
 D. Bravo⁹, B. Caccianiga⁵, F. Calaprice¹⁰, M. Carlini⁸,
 P. Cavalcante^{8,9}, A. Chepurinov¹¹, K. Choi¹², M. Cribier¹³,
 D. D'Angelo⁵, S. Davini², A. Derbin¹⁴, L. Di Noto¹, I. Drachnev²,
 M. Durero¹³, A. Etenko⁴, S. Farinon¹, V. Fischer¹³, K. Fomenko¹⁵,
 D. Franco¹⁶, F. Gabriele⁸, J. Gaffiot¹³, C. Galbiati¹⁰,
 M. Gschwender¹⁷, C. Ghiano¹, M. Giammarchi⁵, M. Goeger-Neff³,
 A. Goretti¹⁰, M. Gromov¹¹, C. Hagner⁷, Th. Houdy^{13,16},
 E. Hungerford¹⁸, Aldo Ianni⁸, Andrea Ianni¹⁰, N. Jonquères¹⁹,
 A. Jany²⁰, K. Jedrzejczak²⁰, D. Jeschke³, V. Kobychov²¹,
 D. Korablev¹⁵, G. Korga⁸, V. Kornoukhov²², D. Kryn¹⁶,
 T. Lachenmaier¹⁷, T. Lasserre¹³, M. Laubenstein⁸, B. Lehnert²³,
 J. Link⁹, E. Litvinovich^{4,24}, F. Lombardi⁸, P. Lombardi⁵,
 L. Ludhova^{25,26}, G. Lukyanchenko^{4,24}, I. Machulin^{4,24}, S. Manecki^{9,27},
 W. Maneschg²⁸, G. Manuzio¹, S. Marcocci², J. Maricic¹²,
 G. Mention¹³, E. Meroni⁵, M. Meyer⁷, L. Miramonti⁵,
 M. Misiaszek^{20,8}, M. Montuschi²⁹, P. Mosteiro¹⁰, V. Muratova¹⁴,
 R. Musenich¹, B. Neumair³, L. Oberauer³, M. Obolensky¹⁶,
 F. Ortica³⁰, M. Pallavicini¹, L. Papp³, A. Pocar³¹, G. Ranucci⁵,
 A. Razeto⁸, A. Re⁵, A. Romani³⁰, R. Roncin^{8,16}, N. Rossi⁸,
 S. Schönert³, L. Scola¹³, D. Semenov¹⁴, M. Skorokhvatov^{4,24},
 O. Smirnov¹⁵, A. Sotnikov¹⁵, S. Sukhotin⁴, Y. Suvorov^{32,4},
 R. Tartaglia⁸, G. Testera¹, J. Thurn²³, M. Toropova⁴, E. Unzhakov¹⁴,
 C. Veyssière¹³, A. Vishneva¹⁵, M. Vivier¹³, R.B. Vogelaar⁹,
 F. von Feilitzsch³, H. Wang³², S. Weinz³³, J. Winter³³, M. Wojcik²⁰,
 M. Wurm³³, Z. Yokley⁹, O. Zaimidoroga¹⁵, S. Zavatarelli¹, K. Zuber²³
 and G. Zuzel²⁰

¹Dipartimento di Fisica, Università degli Studi e INFN, Genova 16146, Italy

² Gran Sasso Science Institute (INFN), 67100 L'Aquila, Italy

³Physik-Department and Excellence Cluster Universe, Technische Universität München, 85748 Garching, Germany

⁴NRC Kurchatov Institute, 123182 Moscow, Russia

⁵Dipartimento di Fisica, Università degli Studi e INFN, 20133 Milano, Italy

⁶Chemical Engineering Department, Princeton University, Princeton, NJ 08544, USA

⁷Institut für Experimentalphysik, Universität, 22761 Hamburg, Germany

⁸INFN Laboratori Nazionali del Gran Sasso, 67010 Assergi (AQ), Italy

⁹Physics Department, Virginia Polytechnic Institute and State University, Blacksburg, VA 24061, USA

¹⁰Physics Department, Princeton University, Princeton, NJ 08544, USA



- ¹¹ Lomonosov Moscow State University Skobeltsyn Institute of Nuclear Physics, 119234 Moscow, Russia
- ¹² Department of Physics and Astronomy, University of Hawaii, Honolulu, HI 96822, USA
- ¹³ Commissariat à l'Énergie Atomique et aux Énergies Alternatives, Centre de Saclay, IRFU, 91191 Gif-sur-Yvette, France
- ¹⁴ St. Petersburg Nuclear Physics Institute NRC Kurchatov Institute, 188350 Gatchina, Russia
- ¹⁵ Joint Institute for Nuclear Research, 141980 Dubna, Russia
- ¹⁶ AstroParticule et Cosmologie, Université Paris Diderot, CNRS/IN2P3, CEA/IRFU, Observatoire de Paris, Sorbonne Paris Cité, 75205 Paris Cedex 13, France
- ¹⁷ Kepler Center for Astro and Particle Physics, Universität Tübingen, 72076 Tübingen, Germany
- ¹⁸ Department of Physics, University of Houston, Houston, TX 77204, USA
- ¹⁹ Commissariat à l'Énergie Atomique et aux Énergies Alternatives, Centre de Saclay, DEN/DM2S/SEMT/BCCR, 91191 Gif-sur-Yvette, France
- ²⁰ M. Smoluchowski Institute of Physics, Jagiellonian University, 30059 Krakow, Poland
- ²¹ Kiev Institute for Nuclear Research, 06380 Kiev, Ukraine
- ²² Institute for Theoretical and Experimental Physics, 117218 Moscow, Russia
- ²³ Department of Physics, Technische Universität Dresden, 01062 Dresden, Germany
- ²⁴ National Research Nuclear University MEPhI (Moscow Engineering Physics Institute), 115409 Moscow, Russia
- ²⁵ IKP-2 Forschungszentrum Jülich, 52428 Jülich, Germany
- ²⁶ RWTH Aachen University, 52062 Aachen, Germany
- ²⁷ Physics Department, Queen's University, Kingston ON K7L 3N6, Canada
- ²⁸ Max-Planck-Institut für Kernphysik, 69117 Heidelberg, Germany
- ²⁹ Dipartimento di Fisica e Scienze della Terra Università degli Studi di Ferrara e INFN, Via Saragat 1-44122, Ferrara, Italy
- ³⁰ Dipartimento di Chimica, Biologia e Biotecnologie, Università e INFN, 06123 Perugia, Italy
- ³¹ Amherst Center for Fundamental Interactions and Physics Department, University of Massachusetts, Amherst, MA 01003, USA
- ³² Physics and Astronomy Department, University of California Los Angeles (UCLA), Los Angeles, California 90095, USA
- ³³ Institute of Physics and Excellence Cluster PRISMA, Johannes Gutenberg-Universität Mainz, 55099 Mainz, Germany

Abstract. The aim of the SOX experiment is to test the hypothesis of existence of light sterile neutrinos through a short baseline experiment. Electron antineutrinos will be produced by a high activity source and detected in the Borexino experiment. Both an oscillometry approach and a conventional disappearance analysis will be performed and, if combined, SOX will be able to investigate most of the anomaly region at 95% c.l. This paper focuses on the improvements performed on the simulation code and on the techniques (calibrations) used to validate the results.

1. Introduction

The SOX experiment aims to investigate the existence of light sterile neutrinos by testing the Gallex and reactor anomalies with a short baseline oscillation experiment [1, 2]. The successful detection of a signal incompatible with the Standard Model prediction could imply the existence of at least one sterile neutrino and consequently it could be a breakthrough in the search for new physics.

The SOX project consists in deploying a high activity ^{144}Ce - ^{144}Pr antineutrino source below the Borexino detector, exploiting the well proven capability of Borexino of detecting antineutrinos via inverse beta decay [3, 4, 5]. We will study the eventual disappearance of antineutrino through both the comparison of the measured and expected antineutrino fluxes and with an oscillometric analysis. The disappearance analysis will rely on the measurement of source activity that will be done using two independent calorimeters [6].

2. Calibration and data analysis

A reliable simulation algorithm of the experiment is fundamental for the data analysis, both for sensitivity studies and for the calculation of the efficiency in detecting the inverse beta decay in the peripheral regions, close to the vessel surrounding the innermost active zone. The Borexino simulation code, which was already used in solar analysis [7], has been extended, in order to be ready for the SOX data analysis. Finally, a novel efficient method for simulating external background events surviving the Borexino passive shield has been developed. It is based on Geant4 toolkit [8]. This technique allows to reliably predict the effect of the external contaminations and can be very useful for the next generation of liquid scintillator experiments. It has been verified comparing the simulations with data acquired with a ^{232}Th source placed in the peripheral part of the detector. The expected source induced background has been evaluated with this method and the result has been compared to those reported in [9], showing good agreement. In the last years, Borexino has been calibrated deploying several vials containing γ sources of different energies in different positions inside the scintillator. These sources have been used to test the accuracy of the simulations in reproducing the collected scintillation light. Moreover, a $^{241}\text{Am} - ^9\text{Be}$ neutron source have been deployed to characterize neutron capture processes which are of fundamental importance for geoneutrino and SOX analyses [10]. A new extensive calibration campaign is foreseen in 2017 before the SOX source arrival.

3. Conclusions

The Borexino detector is an ideal tool to search for sterile neutrinos. It is possible to test the existence of sterile neutrinos using an antineutrino source with almost negligible background. Figure 1 shows the expected SOX sensitivity obtained using the simulations. Uncertainties on the shape factor (σ_b , in units of electron mass) of the beta decay and on the source activity (σ_h) are taken into account. SOX will cover most of the anomaly region (90% confidence level) with the combined analysis (at 95% confidence level). The start of data taking is foreseen in 2017.

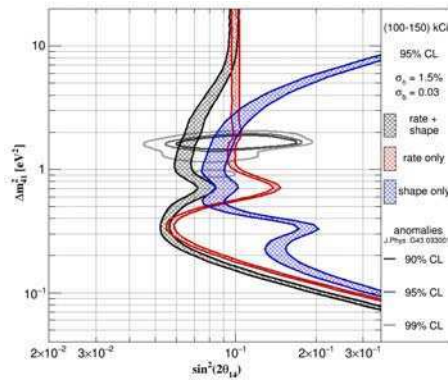


Figure 1. Anomaly region superimposed to the expected SOX sensitivity: rate only (red), shape only (blue) and combined (black).

References

- [1] Mention G *et al.*, 2011, *Phys. Rev. D* **83** 073006
- [2] Giunti C. and Laveder M., 2011, *Phys. Rev. C* **83** 065504
- [3] Bellini G *et al.*, 2013, *JHEP* **38** 1308
- [4] A. Caminata *et al.*, 2016, *Il Nuovo Cimento C* **39** 236
- [5] Agostini M *et al.*, 2015, *Phys. Rev. D* **92** 031101
- [6] L. Papp *et al.*, 2016, *Nuclear Instruments and Methods in Physics Research A* **824** 699700
- [7] Bellini G *et al.*, 2014 *Phys. Rev. D* **89** 112007
- [8] Geant4 Collaboration, Physics Reference Manual
- [9] Gaffiot J. *et al.*, 2015, *Phys. Rev. D* **91** 072005
- [10] Bellini G *et al.*, 2012, *Journal of Instrumentation* **7** 10018