

Towards Equitable, Diverse, and Inclusive science collaborations: The Multimessenger Diversity Network

The IceCube Collaboration and the Multimessenger Diversity Network

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The Multimessenger Diversity Network (MDN), formed in 2018, extends the basic principle of multimessenger astronomy – that working collaboratively with different approaches enhances understanding and enables previously impossible discoveries – to equity, diversity, and inclusion (EDI) in science research collaborations. With support from the National Science Foundation INCLUDES program, the MDN focuses on increasing EDI by sharing knowledge, experiences, training, and resources among representatives from multimessenger science collaborations. Representatives to the MDN become engagement leads in their collaboration, extending the reach of the community of practice. An overview of the MDN structure, lessons learned, and how to join are presented.

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1. Introduction

The lack of diversity in physics, astronomy, and astrophysics is well-documented. For example, 2018 data from the American Institute of Physics' Statistical Research Center shows that only 9% of physics bachelors degrees were earned by Hispanic American students and 3% were earned by African American students [1], and 78% of physics bachelor degrees were earned by men while 22% were earned by women [2]. Increasing diversity in the fields of physics and astronomy is important to ensure scientific progress by making sure all talent is being tapped, in addition to the ethical and social justice motivation. This is especially true in large scientific collaborations which increasingly play a primary role in a researcher's professional interactions and research opportunities.

The Multimessenger Diversity Network (MDN), formed in late 2018 with four founding members, is now a community of practice comprised of representatives from nine multimessenger astronomy and astrophysics experiments. Current participating collaborations include the *Fermi* Gamma-ray Space Telescope, IceCube Neutrino Observatory, Laser Interferometer Space Antenna, LIGO Scientific Collaboration, North American Nanohertz Observatory for Gravitational Waves, Pierre Auger Observatory, Neil Gehrels *Swift* Observatory, Vera C. Rubin Observatory, and the Very Energetic Radiation Imaging Telescope Array System. The initial funding was obtained through a supplemental award to IceCube through the NSF INCLUDES (Inclusion across the Nation of Communities of Learners of Underrepresented Discoverers in Engineering and Science) program which aims to transform education and career pathways to broaden participation in science and engineering. As an INCLUDES-funded program, the MDN focuses on broadening participation in multimessenger astronomy by sharing resources, knowledge, experiences, and training among participating collaborations.

In this proceeding, we offer an overview of the MDN components, highlight initial impacts of the network, share lessons learned, and provide information on how to join the group.

2. MDN Components

The MDN is based on a "community of practice" model [4] introduced to members in 2019 in a community engagement training workshop led by Lou Woodley, Director of the Center for Scientific Collaboration and Community Engagement (www.cscce.org/). A community of practice is a group of people who care about a subject and carry out activities or share resources on the subject of interest. The MDN relies on and promotes six structural elements shown (Figure 1) to advance equity, diversity, and inclusion (EDI) in multimessenger collaborations: (1) *opportunity* to go beyond individual accomplishments, (2) *structure* through organizational principles and tools, (3) *training* for members, (4) *support* from each other and for current and future science, technology, engineering and mathematics (STEM) professionals, (5) *presence* at conferences, on websites, and on media outlets, and (6) *legitimacy* in broadening participation efforts. These elements underpin monthly meetings of the network which provide opportunities for members to share experiences and knowledge, learn from others, and seek and provide support. A website (astromdn.github.io), along with presentations at conferences, and a submitted Astro2020 State of the Profession white paper [5], lend legitimacy to the often volunteer-based EDI efforts of collaborations.

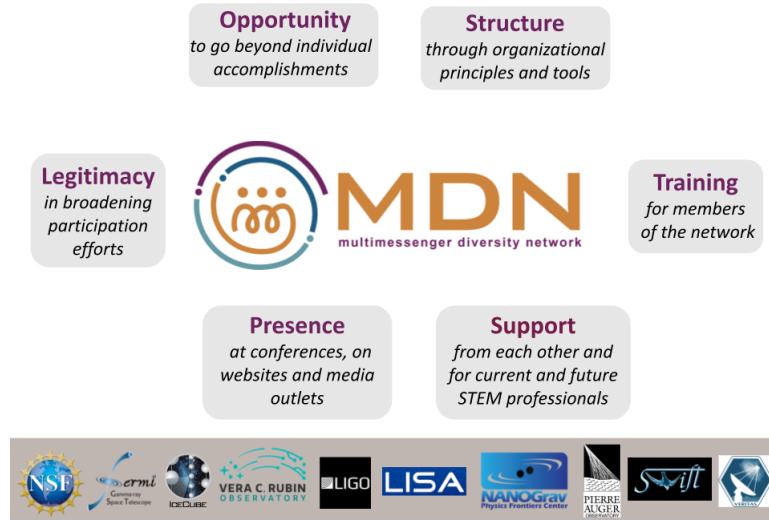


Figure 1: The elements of the MDN and current members, the *Fermi* Gamma-ray Space Telescope, IceCube Collaboration, Vera C. Rubin Observatory, Laser Interferometer Gravitational-Wave Observatory (LIGO), Laser Interferometer Space Antenna, Nanohertz Observatory for Gravitational Waves (NANOGrav), Pierre Auger Observatory, Neil Gehrels *Swift* Observatory, and Very Energetic Radiation Imaging Telescope Array System (VERITAS).

A part-time community manager runs the monthly meetings, responds to inquiries, helps represent the MDN at conferences, and promotes engagement within the group. This position has proven critical to the sustainability and longevity of the network. The community manager keeps the group connected, engaged, and helps rally members around a few key activities that catalyze interactions. In the past, activities included co-authoring an Astro2020 State of the Profession white paper, participating in a two-day community engagement training, and hosting guest speakers. Some planned activities were paused due to the COVID-19 pandemic.

As a community of practice, the MDN focuses on EDI in multimessenger astronomy collaborations with regular activities including monthly meetings, maintaining a website, and participating in conferences. The Community Participation Model from the Center for Scientific Collaboration and Community Engagement, Figure 2, sheds light on stages of community development and has been a useful model for the MDN. As Woodley and Pratt [3] write, communities often start in the "convey/consume" phase with information moving out from a community manager to community members, and move towards a "co-create" phase where members work together to create something new. Reflecting on the MDN as a community, it has occupied each community participation phase at some point, with much time spent advancing and retreating between "collaborate" and "co-create." Referencing Figure 2, the network's interactions, goals, and activities are most aligned with these two phases of community participation.

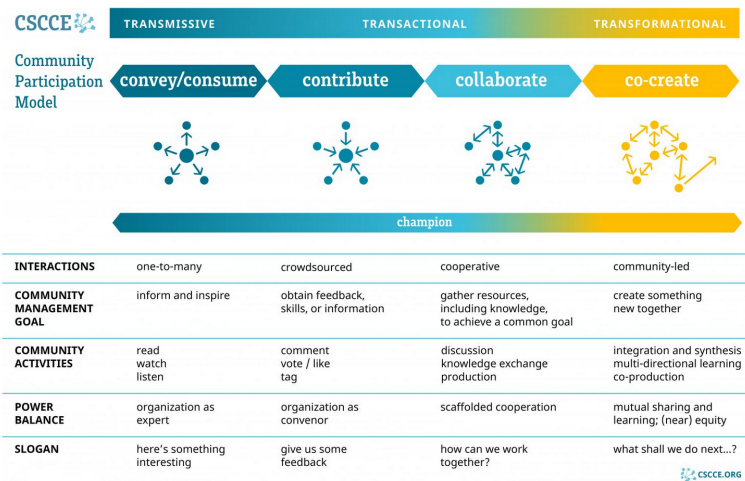


Figure 2: The Community Participation Model from the Center for Scientific Collaboration and Community Engagement [3] describes four modes of community member participation. With both cooperative and community-led interactions, participation in the MDN falls under "collaborate" and "co-create" community models.

3. Initial Impacts

The strength of the MDN lies in the community connections which provide opportunities to share experiences, describe lessons learned, exchange documentation for best practices such as a code of conduct, and present models of a variety of EDI efforts. Raising awareness of EDI work among members is one positive outcome on the MDN effort. Here we highlight three examples exemplifying the community of practice model where members are co-creators and freely exchange ideas and materials for others to adapt and adopt.

The VERITAS Collaboration has moved forward with two major DEI efforts, namely the recognition of collaboration service through annual VERITAS Outstanding Contribution Awards, as well as the adoption of an official Collaboration Code of Conduct. Both of these positive changes were motivated through the communication of procedures already in place within other MDN member collaborations. In the case of the service awards, which are given to early-career scientists in recognition of critical contributions to the collaboration, the IceCube Impact Awards were a model. The VERITAS collaboration Code of Conduct was motivated by and styled after the *Fermi-LAT* collaboration version.

In bringing together several multimessenger collaborations, it is clear that many groups are choosing similar actions to address their EDI disparities. Two topics of interest within the MDN over the last two and half years are collaboration-wide surveys and mentoring programs.

(1) Collaboration-wide surveys: In the course of pursuing EDI initiatives, two questions that often come up are, "How will we know if our efforts are working if we don't have any demographic information about our members?" and "How does the collaboration experience differ by group (broken out by gender, career level, and/or race)?" One way to address one or both of these questions is a collaboration-wide survey. Such a survey could focus on collecting demographic data or focus

on the climate, or atmosphere, of a collaboration. As an example within the MDN, the survey experiences of NANOGrav have directly informed the discussions of a survey within the IceCube collaboration. While the IceCube survey effort is in the very early planning phases, it has been invaluable to hear how another geographically distributed collaboration has developed and evolved their own survey strategy. NANOGrav representatives have shared with some frequency about their collaboration-wide surveys, not just about the planning and execution, but critically about how the collaboration takes action based on the survey results. As a result of recent surveys, NANOGrav has put in place more support structures for post-doctoral scholars and graduate students, and restructured its membership to be more transparent and accessible to early career researchers. Such examples are a reminder that running a survey should not be an end goal. They are one way to collect data on the pulse of a collaboration and help set a course for meaningful and sustainable change. Sharing the lessons from NANOGrav with the other eight collaborations shows the value of the MDN to leverage the effort of one group to benefit a much larger number of researchers and scholars.

(2) Mentoring programs: In the case of mentorship, a recent experience highlights the potential of having a community of practice focused on EDI in multimessenger astronomy. The suggestion for a meeting on mentorship came from the *Fermi*-LAT representative, who connected the MDN community manager with the mentoring program lead. Members of the MDN were encouraged to invite collaborators to join and several new individuals were welcomed on the call. After a presentation on the *Fermi*-LAT graduate student mentoring program, the group asked questions and engaged in discussion. The *Fermi*-LAT group graciously shared their resources with the MDN including a program proposal and interest forms, and encouraged others to use them as springboards for mentoring programs in their own collaborations. This had a direct impact on the LISA Consortium, which is now pursuing a mentoring program. This generous sharing of knowledge and experience is at the core of the MDN and amplifies the work of a few across the field at-large.

4. Lessons Learned

Since the MDN was formed in 2018, EDI in STEM have taken on a new level of recognition and importance. Now more than ever, researchers, scientists, and scholars are engaged in work to make STEM a more equitable, diverse, inclusive, and accessible field. Here, we share lessons learned from the last two and half years.

Despite the growth in awareness and importance of EDI in STEM, this work is still overwhelmingly volunteer-based, i.e. not necessarily counted in the work load. It is challenging for individuals, many of whom were already addressing work-life balance issues and are often from underrepresented groups, to dedicate time to EDI efforts. This reality has implications for how people participate in the MDN. For instance, a regular meeting cadence is sometimes the only opportunity for engagement during the month and has become the key vehicle for knowledge sharing and community building. Additionally, planned activities have been adapted or abandoned based on the interest and availability of members. At the start of the MDN, it was thought that members would develop a "playbook" [6] or "guidebook" of EDI in their collaboration. The idea and documentation were proposed and included in a training event. Although members recognized that such a guide-

book could be useful, few had the time or energy to invest in its development. The network pivoted to rely on knowledge sharing *among* collaborations during the monthly meetings and through a dedicated Slack organization. For now, this level of engagement is sufficient to support members as they work for organizational change with the time and energy they have available. We continue to push for a recognition of EDI work as critical for the well-being and advancement of scientific collaborations. Counting EDI work towards collaboration-level service, highlighting EDI work in press releases, and developing EDI-focused leadership roles are actionable steps collaborations could take to recognize this important effort [5].

5. Joining MDN

When the MDN formed, it was unclear how much interest there would be in such a community. Through contacts and professional networks, the MDN grew quickly from four to six to nine member groups; there are regular expressions of interest from the broader fields of high-energy, astroparticle physics, and cosmology. If you are interested in joining or following our work, please visit astromdn.github.io and subscribe to our Google Group, or write directly to the contact authors. We welcome your participation and interest.

6. Summary and Outlook

STEM professionals continue to be overwhelmingly male and do not reflect the overall racial and ethnic makeup of the world as a whole. We described the MDN, a community of practice with representatives from nine larger multimessenger astronomy and astrophysics experiments, which is working to increase awareness of EDI issues and achieve better outcomes in the multimessenger community. The elements of the MDN and its structure were described as well as the central role of the community manager who acts as a facilitator to promote interactions and exchanges within the network. A few of the lessons learned by the MDN were introduced and an invitation with instructions on how to get more information was presented.

The outlook for the MDN is encouraging, with a core group that meets virtually monthly in a supportive, educational environment. The EDI challenges are significant and, as described, often being addressed on top of other expectations that usually have higher priority. The MDN promotes the value of EDI to both achieve better outcomes and also the recognition of the need and value of this work.

References

- [1] American Institute of Physics, *The Proportion of Physics Bachelor's Degrees Awarded to African Americans and Hispanic Americans*.
www.aip.org/statistics/data-graphics/proportion-physics-bachelors-degrees-awarded-african-americans-and-0
- [2] Mulvey, P. J. and Nicholson, S. (August, 2020), *Physics Bachelor's Degrees: 2018. Results from the 2018 Survey of Enrollments and Degrees*.
www.aip.org/statistics/reports/physics-bachelors-degrees-2018

- [3] Woodley, L. and Pratt, K., *The CSCCE Community Participation Model – A framework for member engagement and information flow in STEM communities.*
zenodo.org/record/3997802#.YNSP9RGS1PZ
- [4] Wenger-Trayner, E. and Wenger-Trayner, B., *Introduction to communities of practice*
wenger-trayner.com/introduction-to-communities-of-practice/
- [5] Bechtol, E., Bechtol, K., BenZvi, S., Cenko, B., Corlies, L., Couvares, P., Furniss, A., Kaplan, D.L., Key, J.S., Madsen, J., McNally, F., McLaughlin, M., Santander, M., Vigeland, S., and Zuniga-Paiz, J., *Astro2020 APC White Paper: Pursing diversity, equity, and inclusion in multimessenger astronomy collaborations over the coming decade*
arxiv.org/abs/1907.06970
- [6] O'Donnell, S., *Building a Community Playbook Part 1: What is it?*
www.cscce.org/2017/05/26/building-a-community-playbook-pt-1-what-is-it/

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